

Reconnection Outflows in the Extended Corona and Magnetotail



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Matthew J. West, Daniel B. Seaton

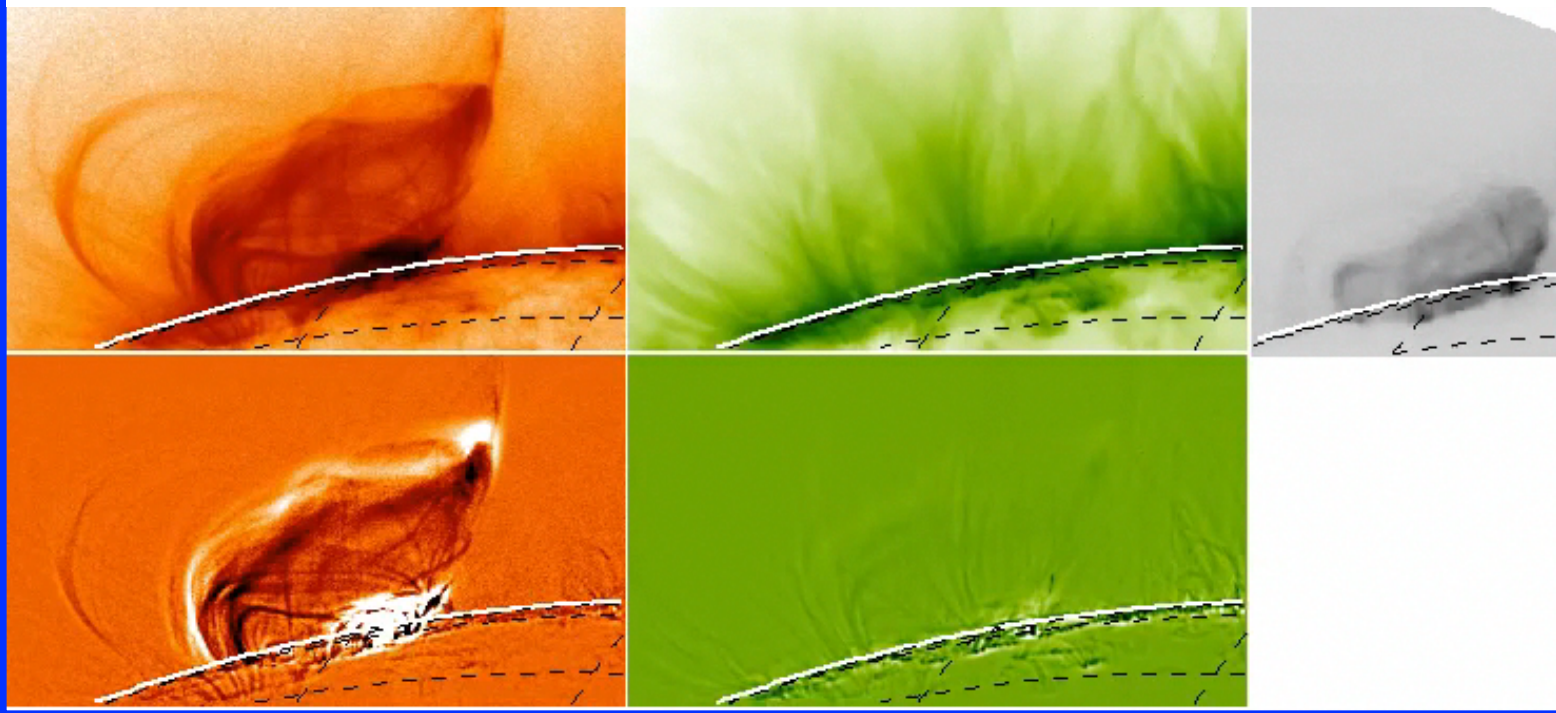


Royal Observatory
of Belgium

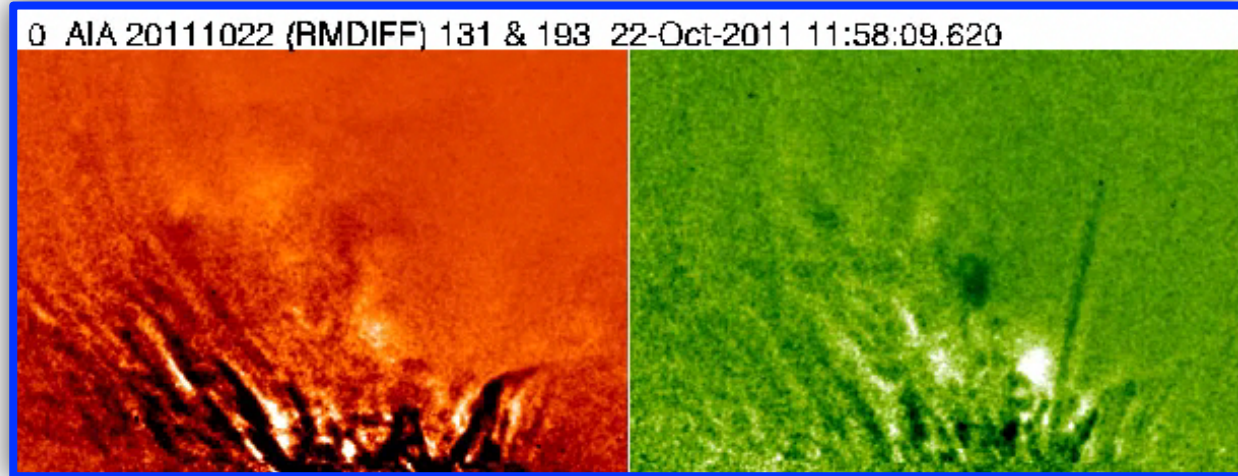
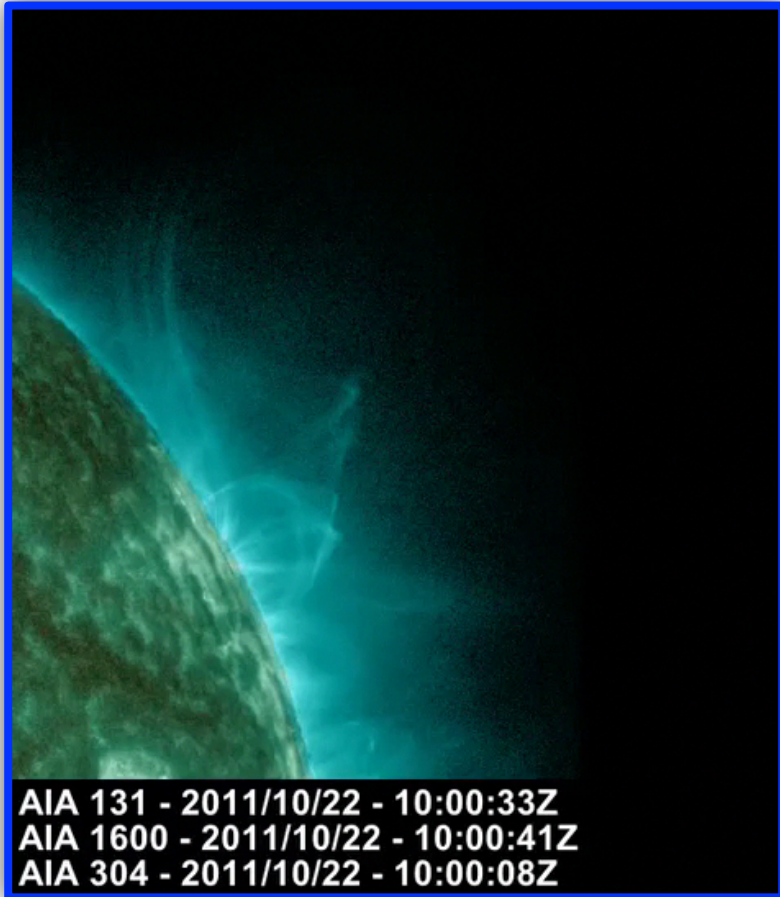


Supra-Arcade Downflows (SADs) and Downflowing Loops (SADLs) Observations

2012 Jan 27, SDO/AIA 131, 193, Hinode/XRT



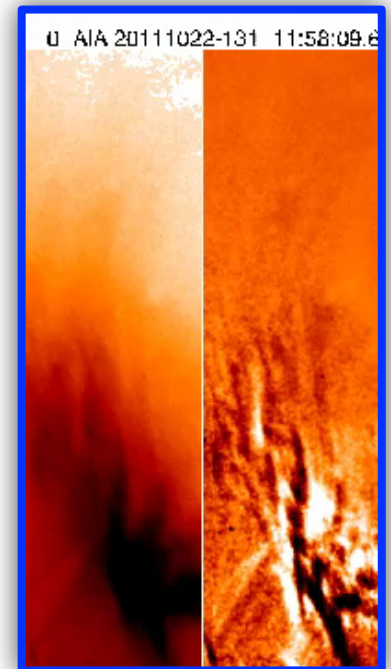
SADs + SADLs



Hot AIA channels. (2-10 MK)

Run-mean-differenced

Reverse-scaled





SADs + SADLs

- Key features
 - Different from plasmoids
 - Observationally associated with inflows (as outflows)
 - Significant correlations with particle acceleration and heating (temporally and spatially), thanks to RHESSI and radio observations

Diagram Models

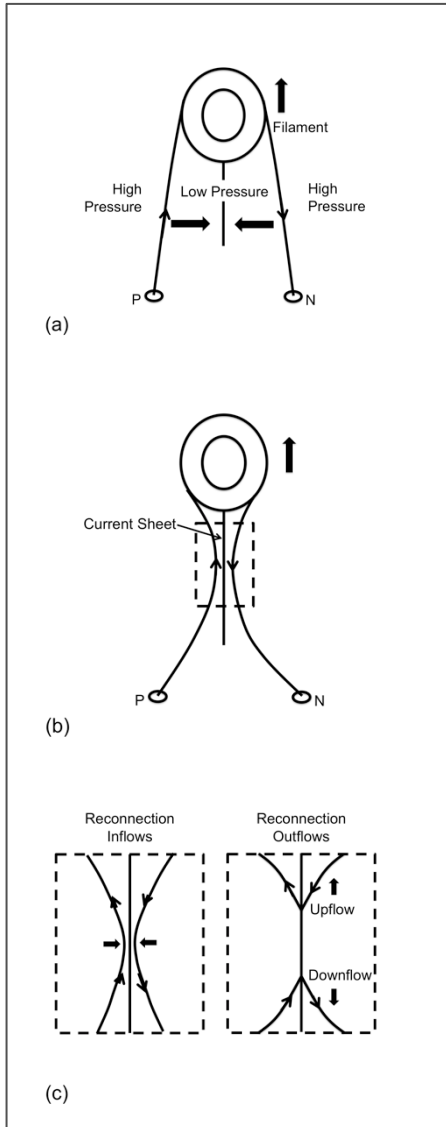


Fig 1

Fig 2

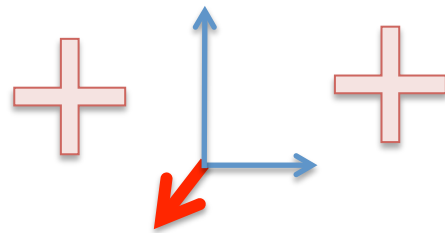
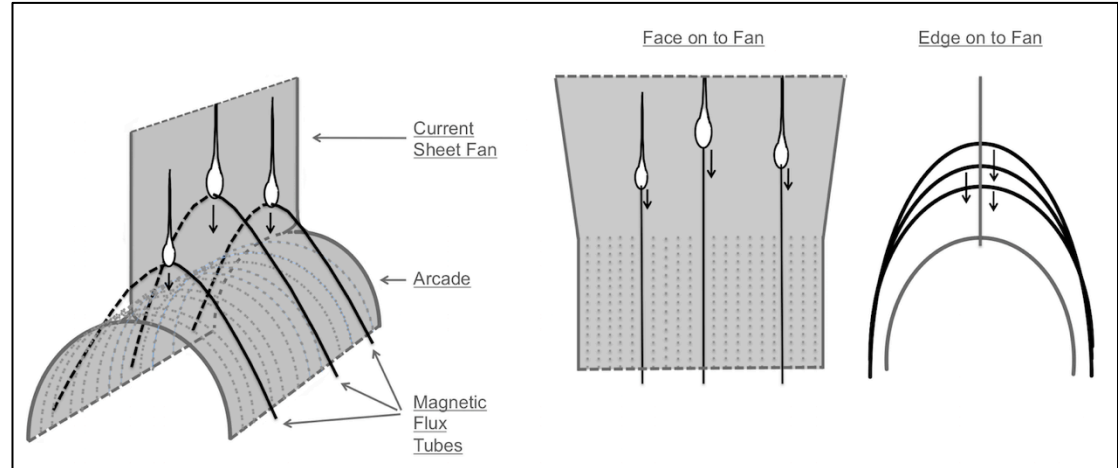


Fig 3

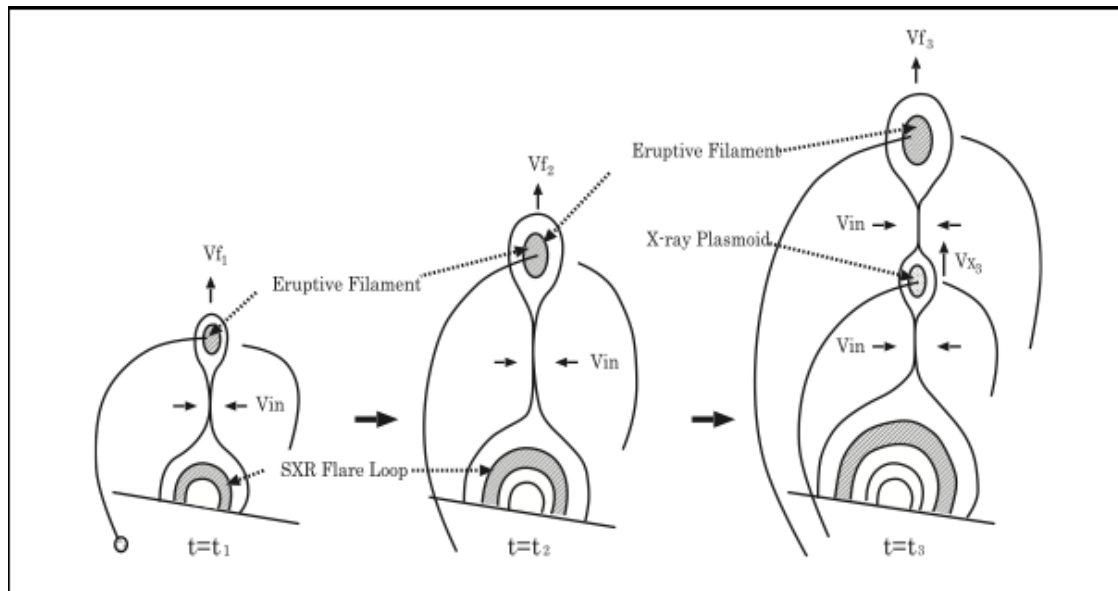
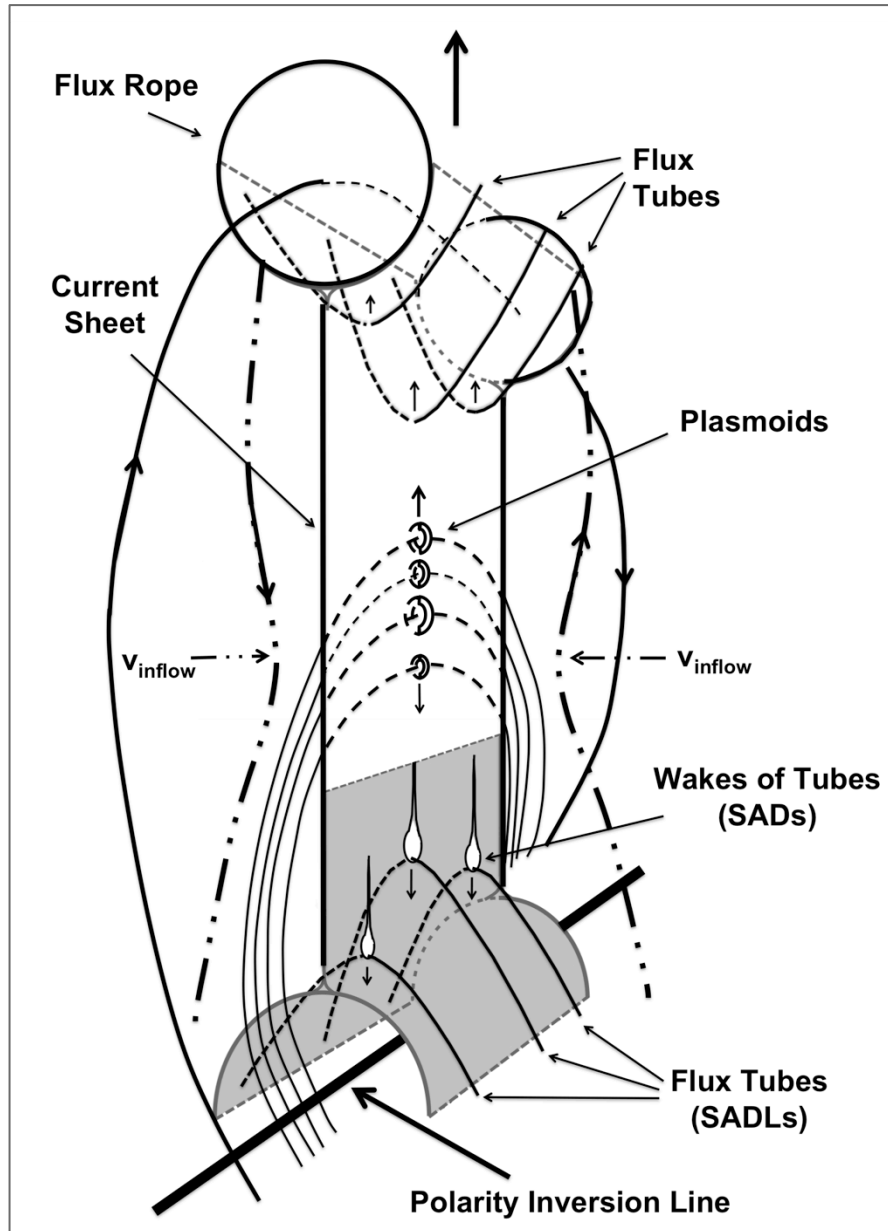


Fig 3: Ohyama & Shibata 2008

Fig 1, 2: Savage et al. 2012

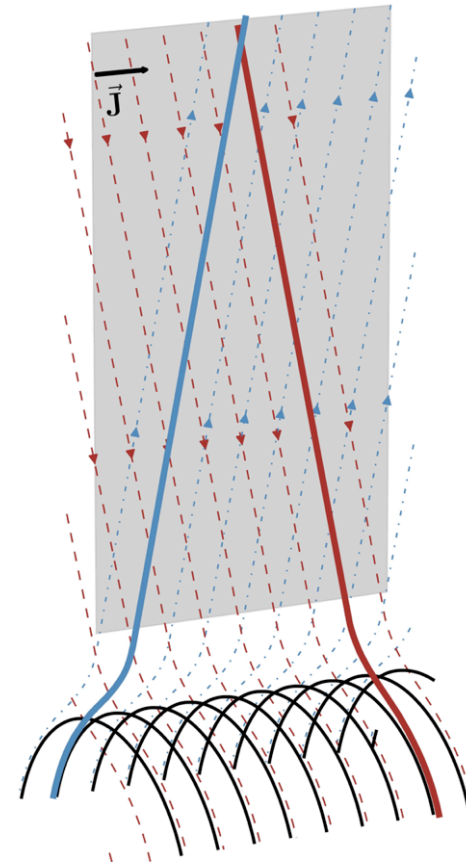


Fig 1



Basic reconnection scenario, post initial flux rope formation and release.

- Field lines reconnect across the current sheet to form outflowing flux tubes while plasmoids form along the current sheet.
- SADs are formed as the flux tubes (SADLs) retract through hot plasma in the fan (*otherwise, only SADLs are observed*).



SADs in the Extended Corona

Fig 1

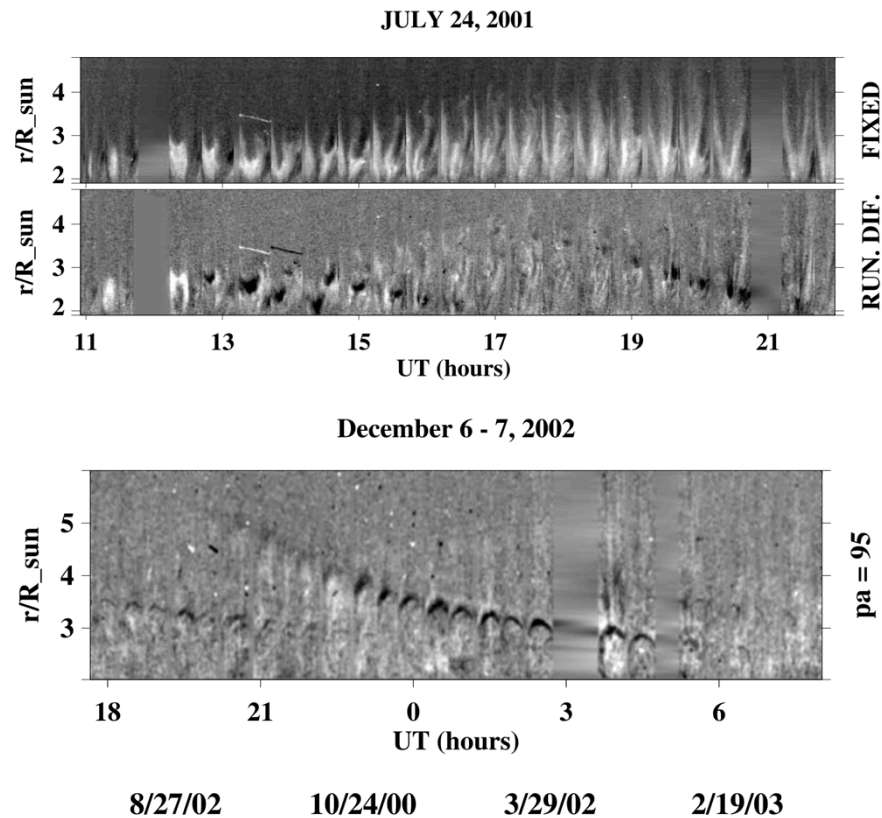
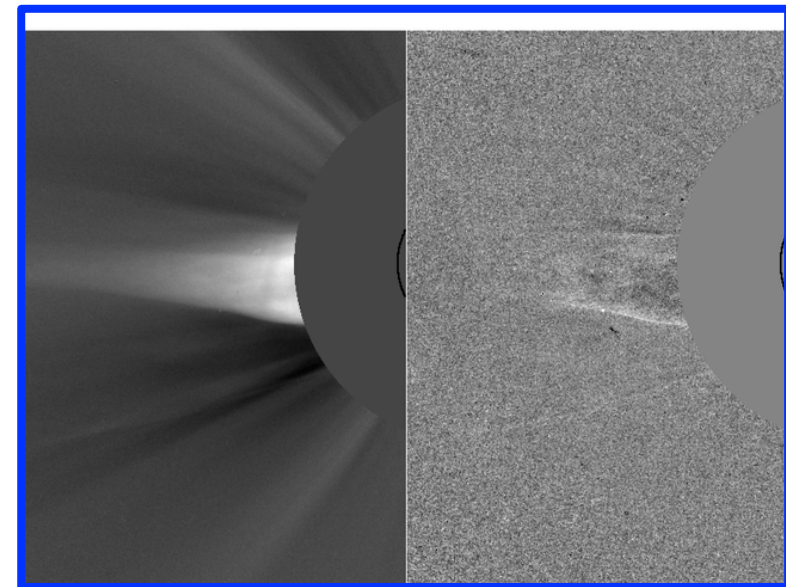
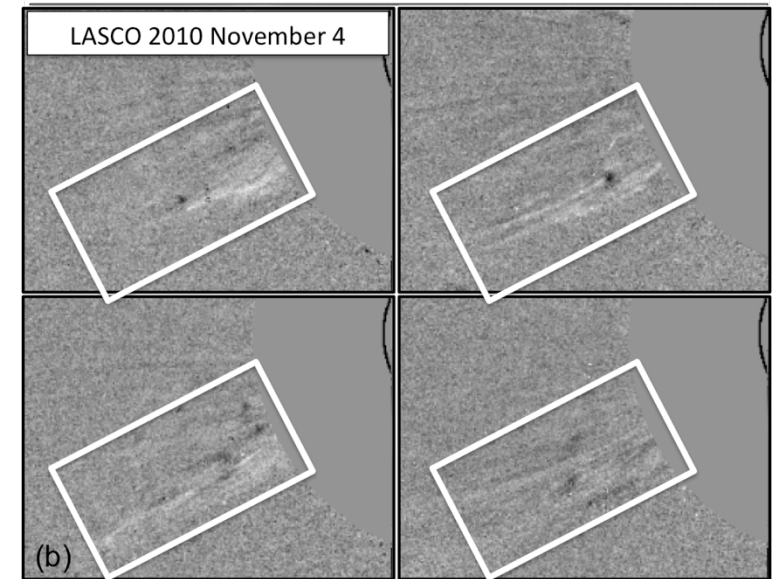
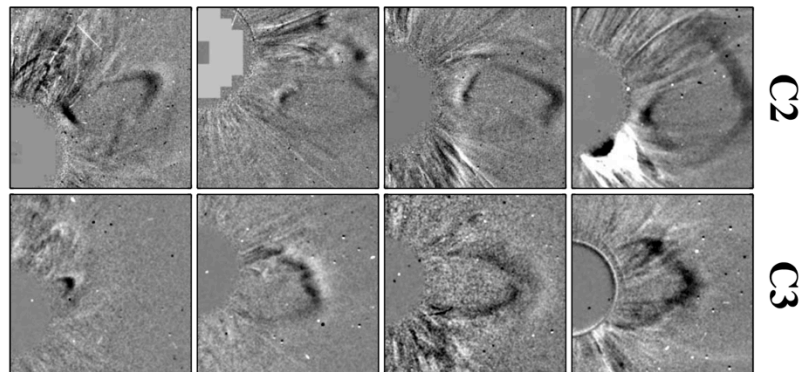


Fig 2

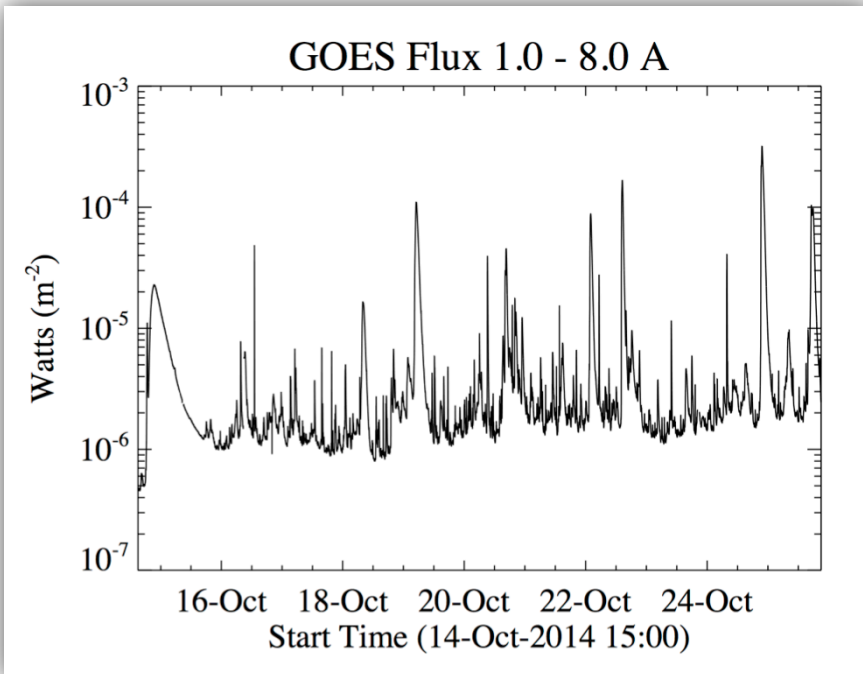


SADs in the Extended Corona...

SADs in the lower corona are typically observed well after reconnection has occurred.

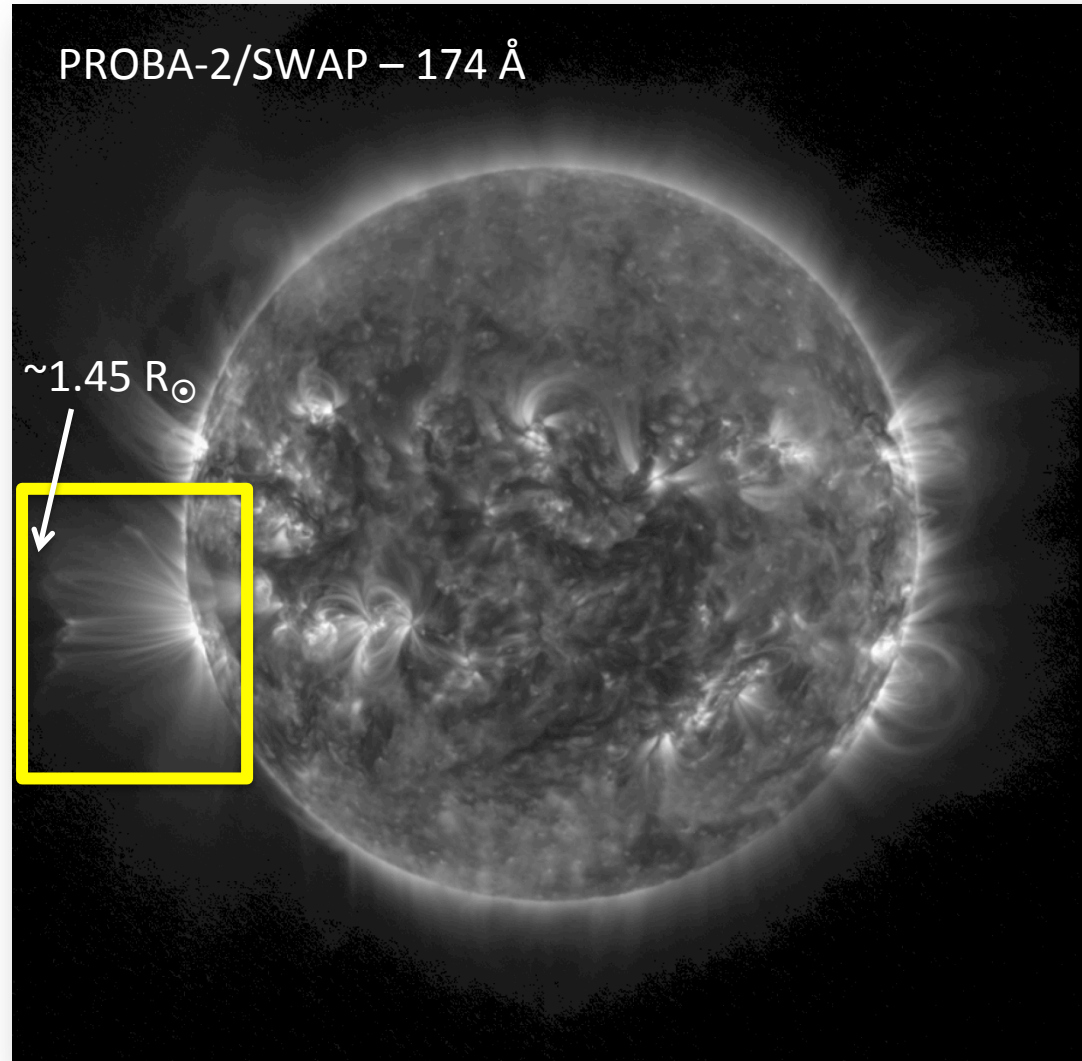
In the extended corona, we are better able to observe the migrating reconnection sites.

WL coronagraphs allow us to see reconnection develop behind the CME while looking directly at the density.



“Giant Arches” Flare – 2014 Oct 14

Fig 1



SADs in the Extended Corona...

LASCO C2

PROBA-2/SWAP

AIA 131 Å

A: Flattened from a year's worth of data

Cleaned (cosmic rays, background stars, planets)

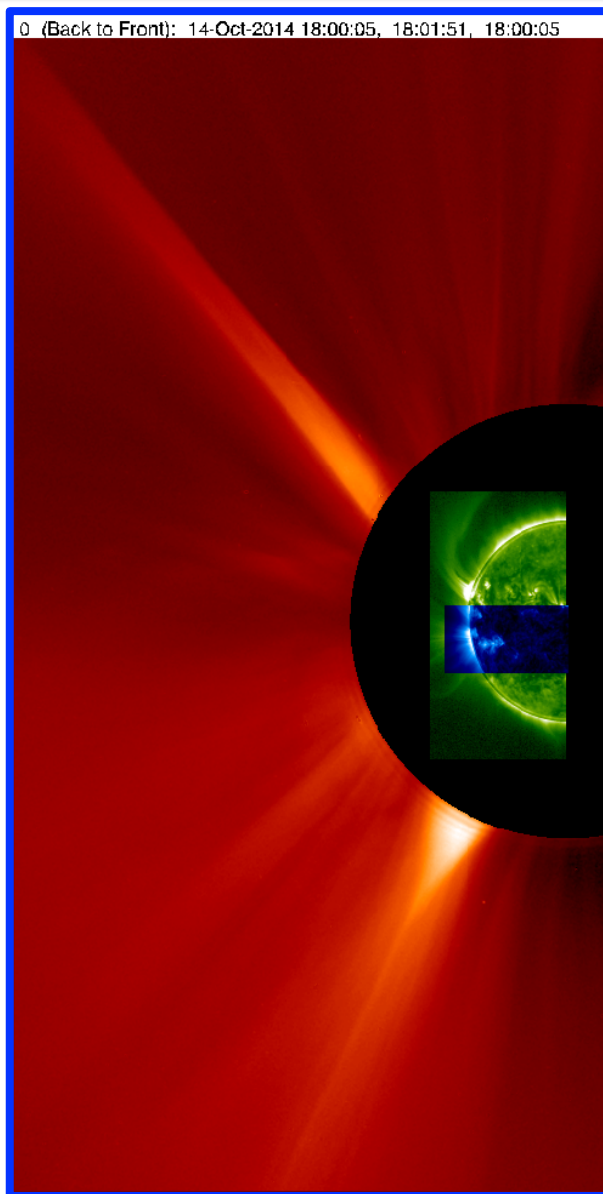
Attenuated disk

B: All that +
Run-mean-differenced

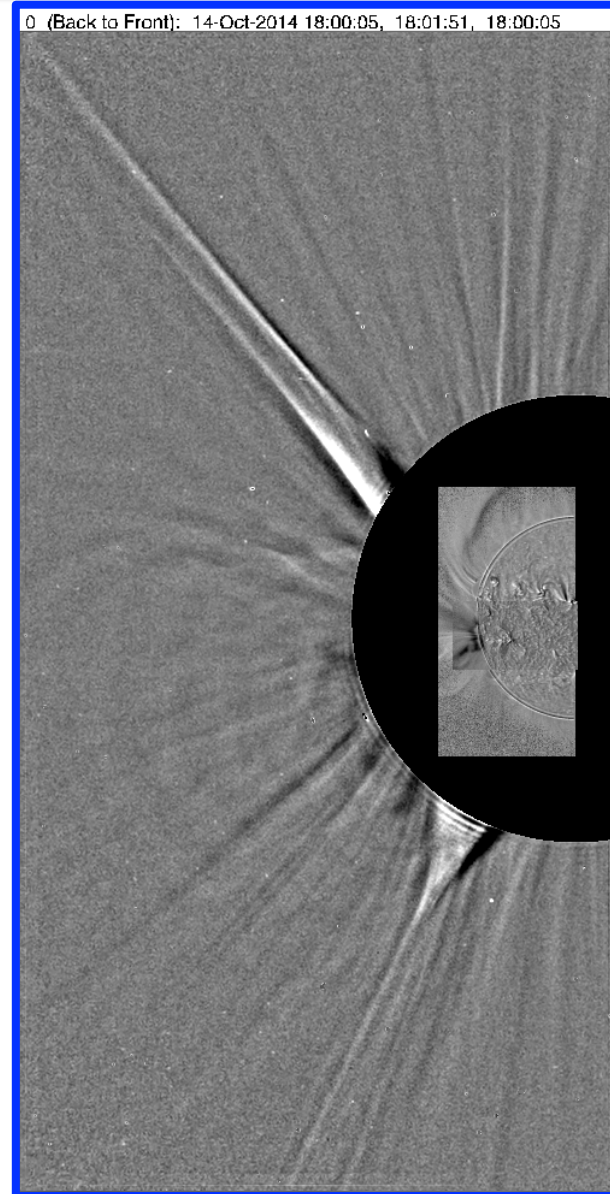
- Both Scaled

Downflows in C3 as well!

A



B



SADs in the Extended Corona...

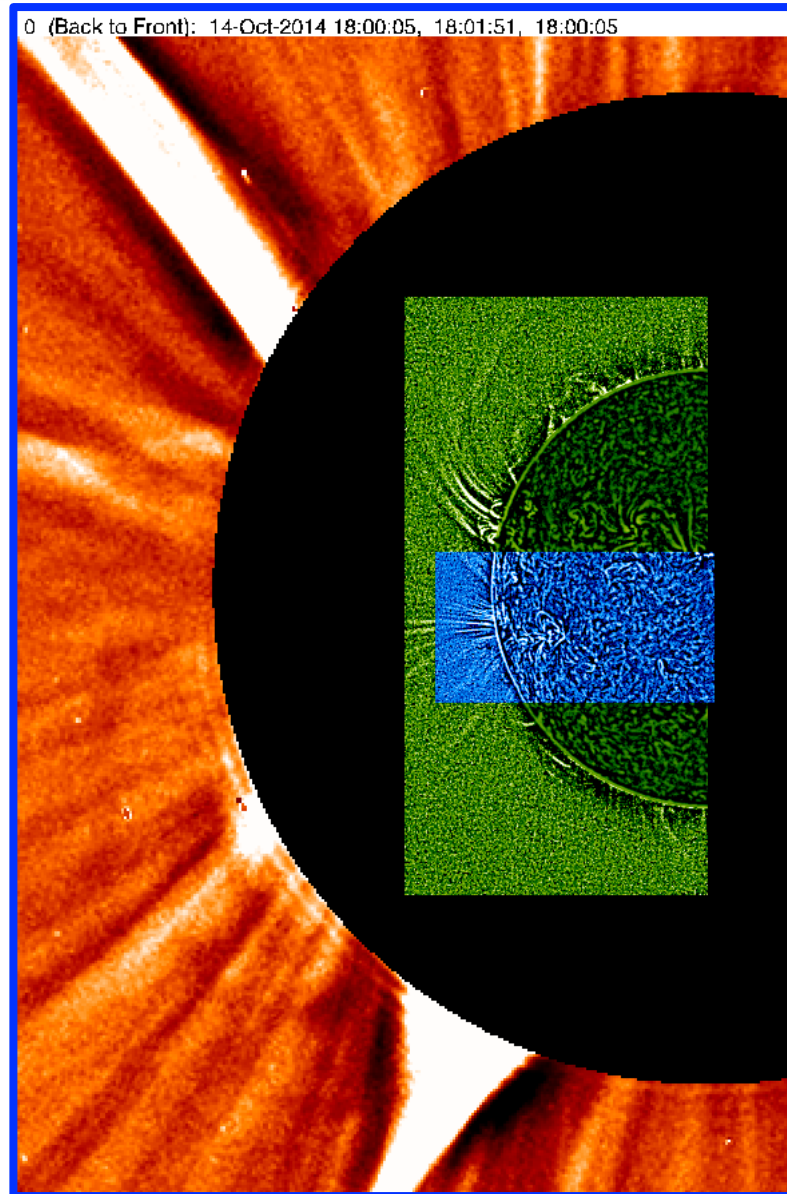
LASCO C2

PROBA-2/SWAP

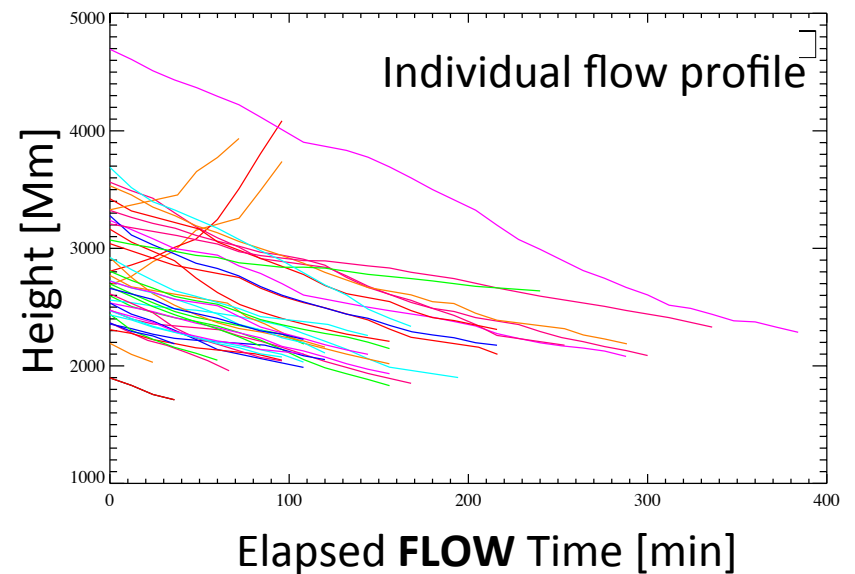
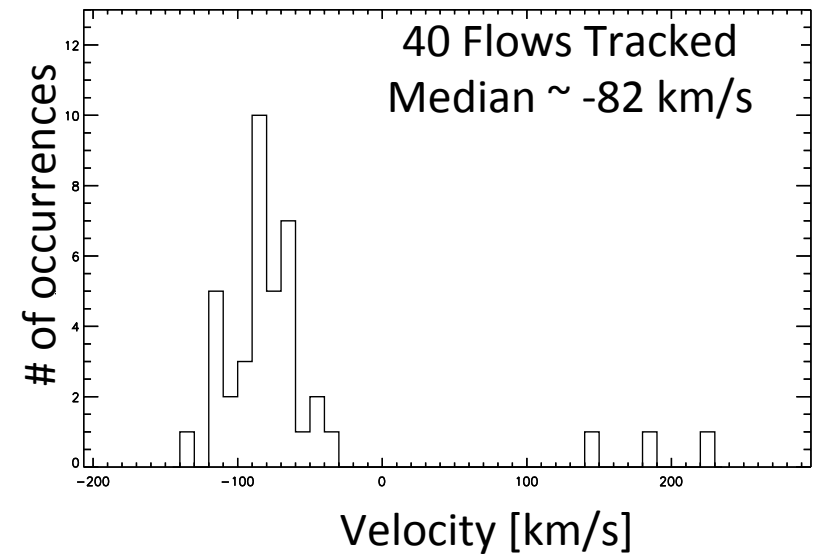
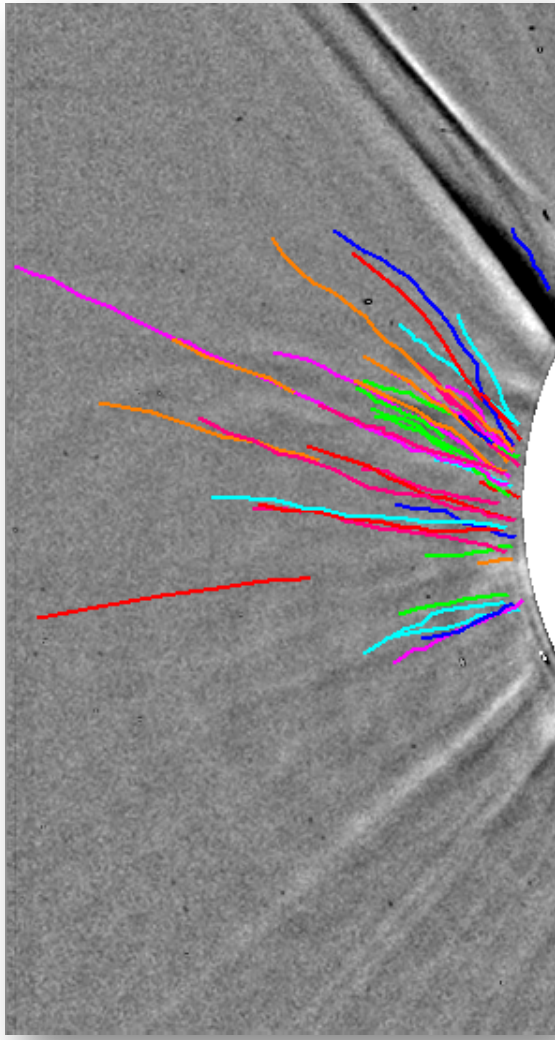
AIA 131 Å

C: Smooth-Differenced
Extracted
Scaled

C



SADs in the Extended Corona...



Strong potential analogy with magnetotail substorms

Fig 1

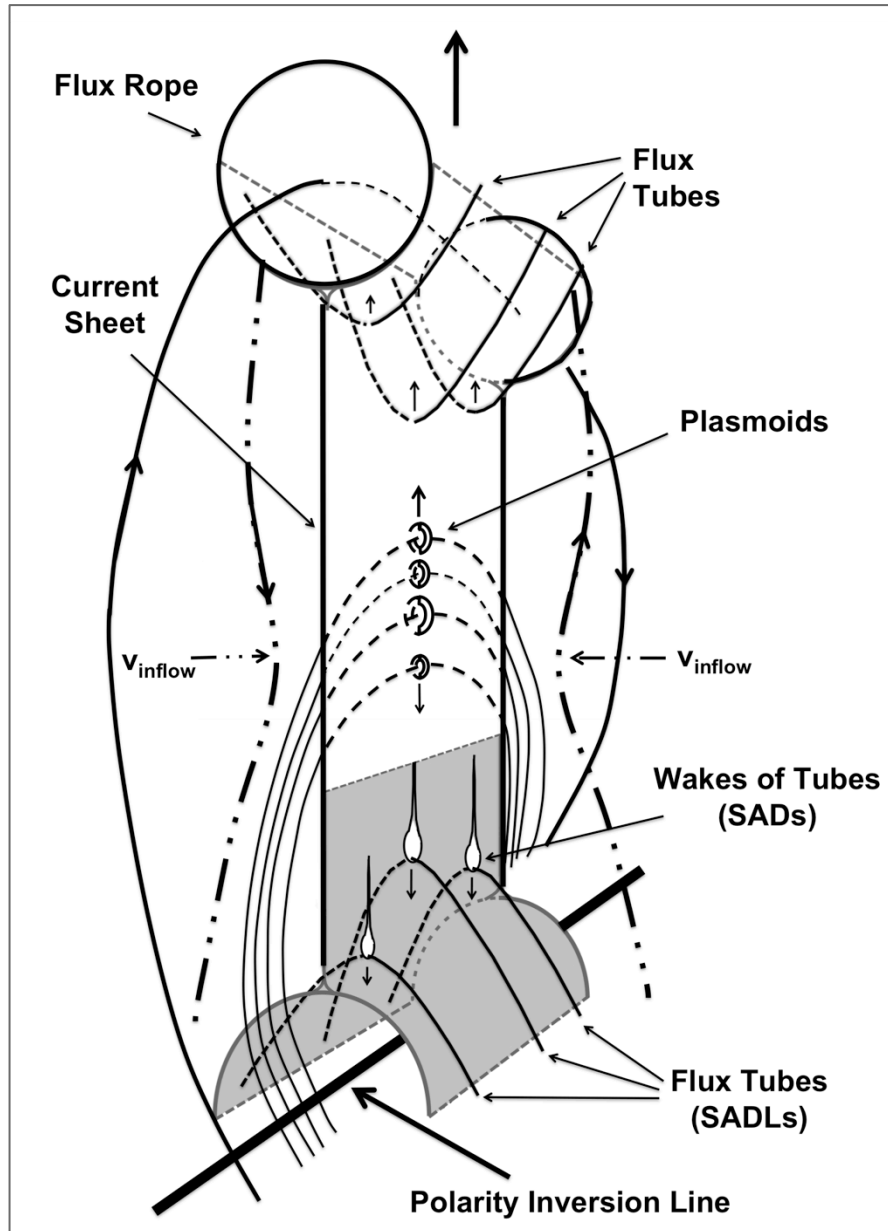
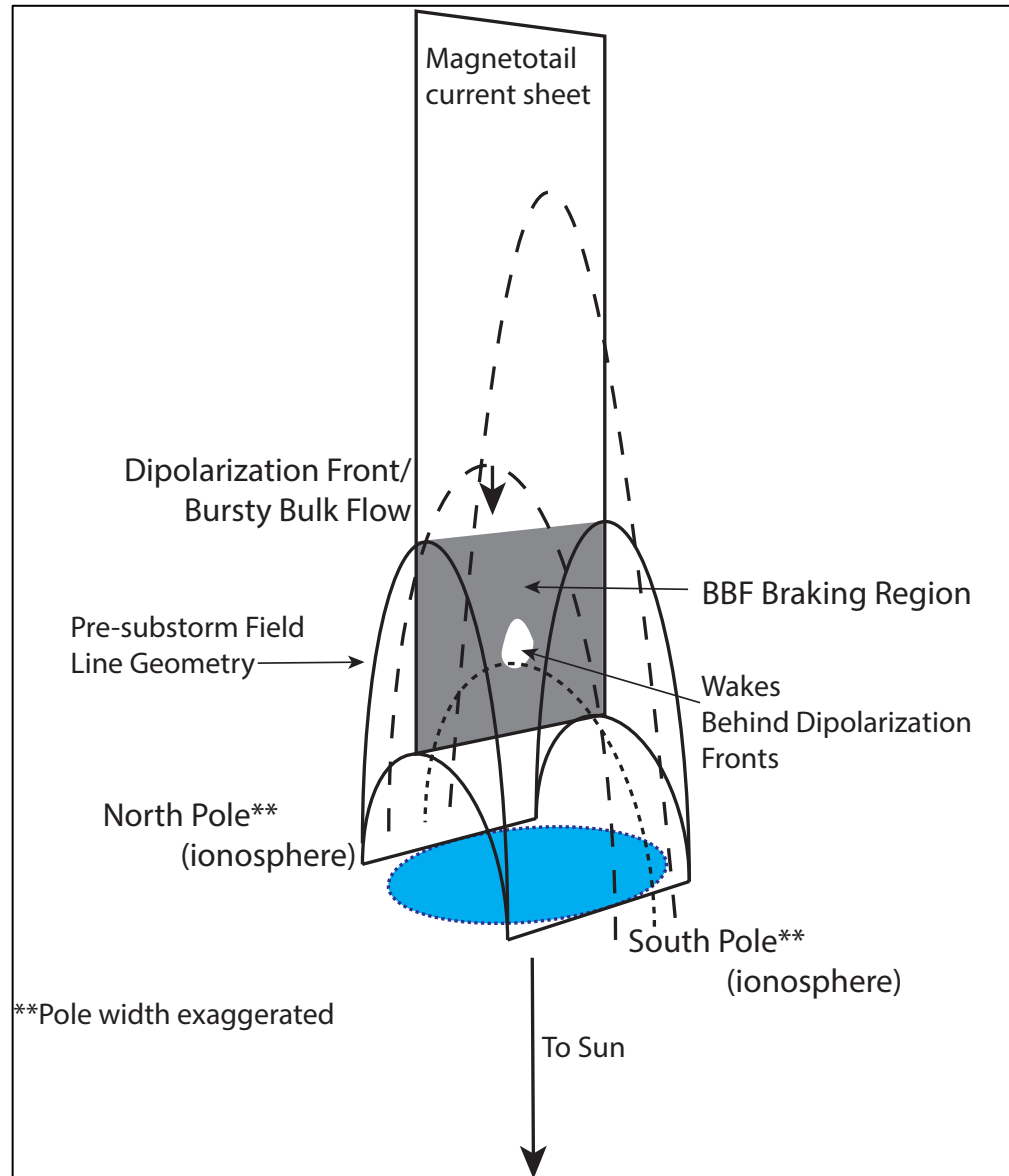
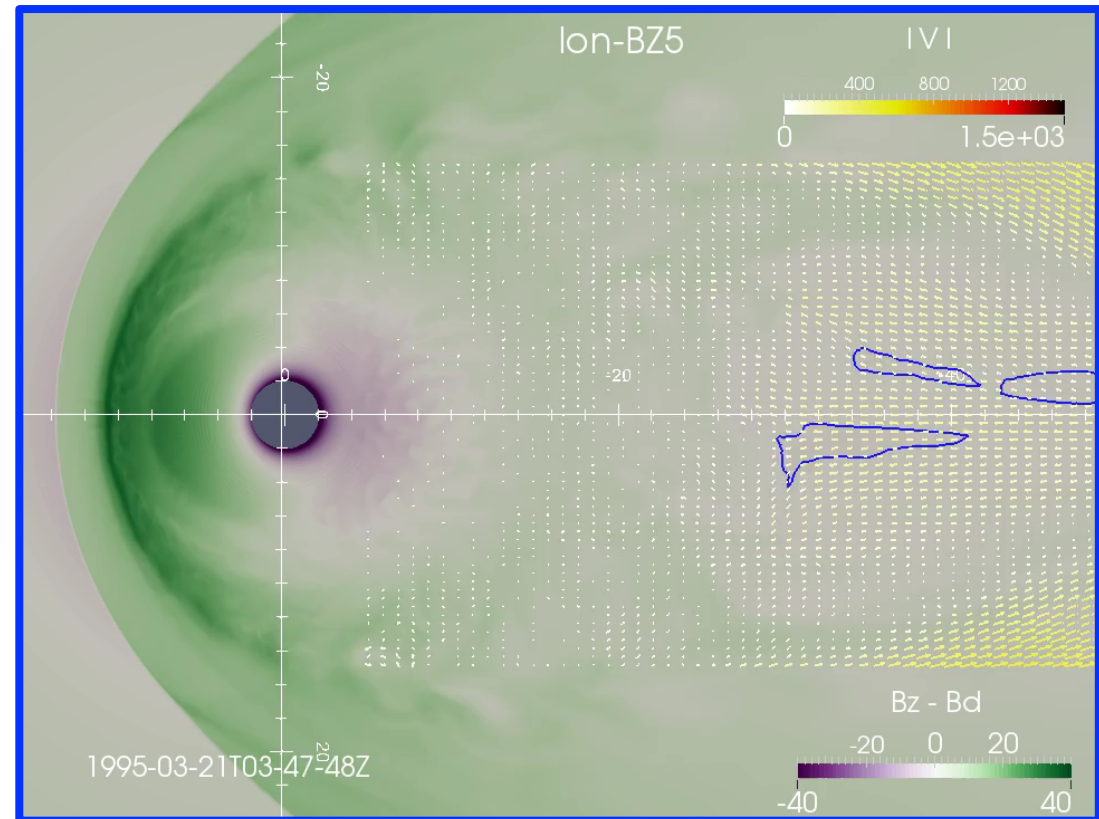
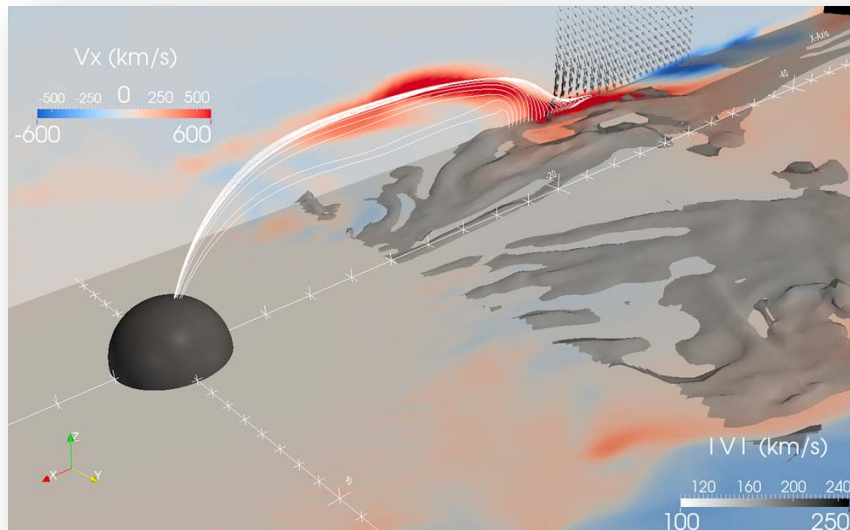


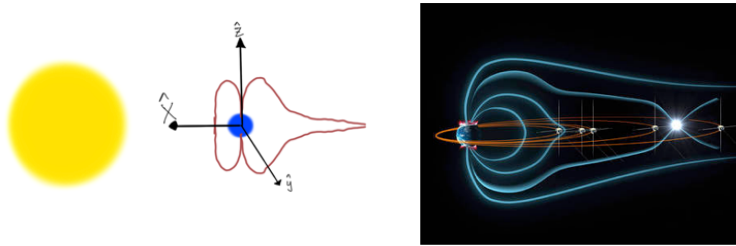
Fig 2



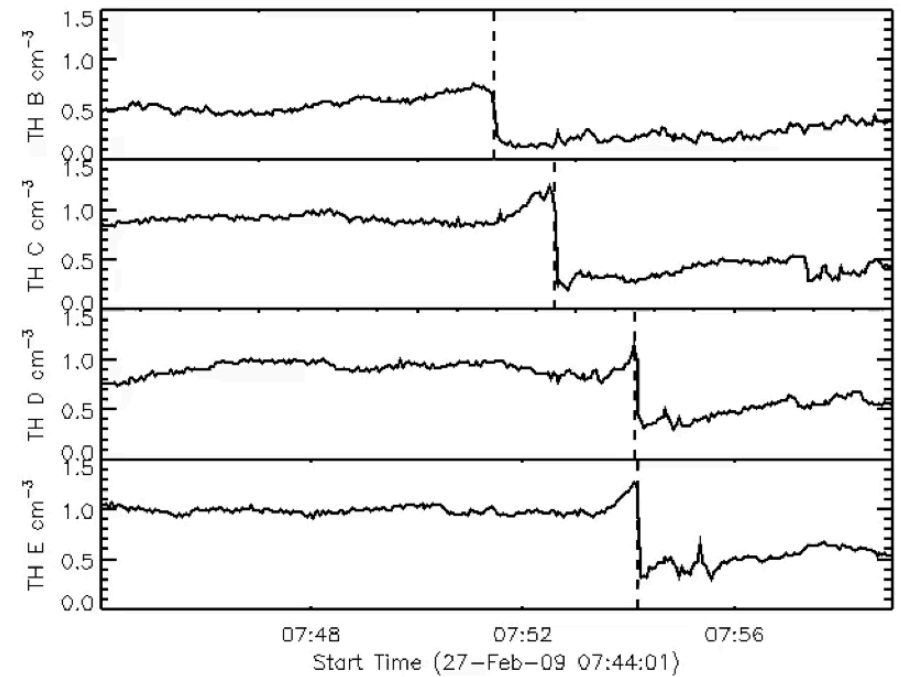
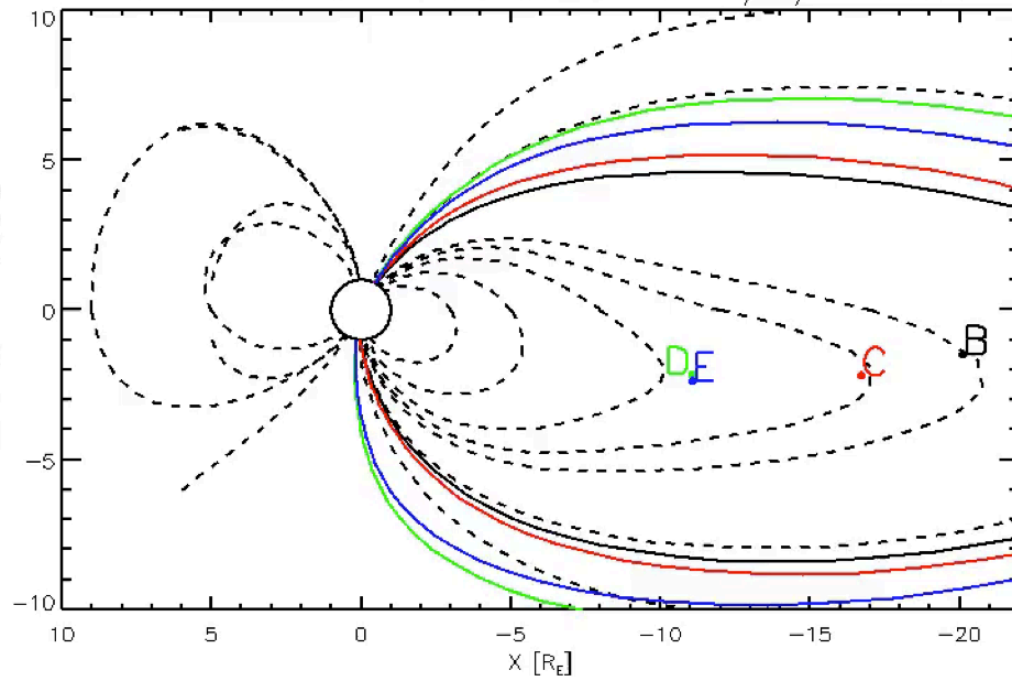
Using THEMIS for comparison to *in situ* data



Using THEMIS for comparison to *in situ* data

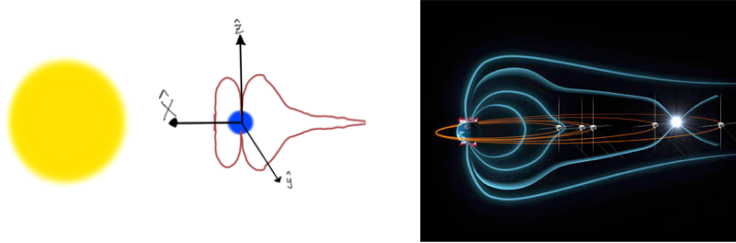


Model t01 and Location of THEMIS Probes 2009/02/27 07:44:00

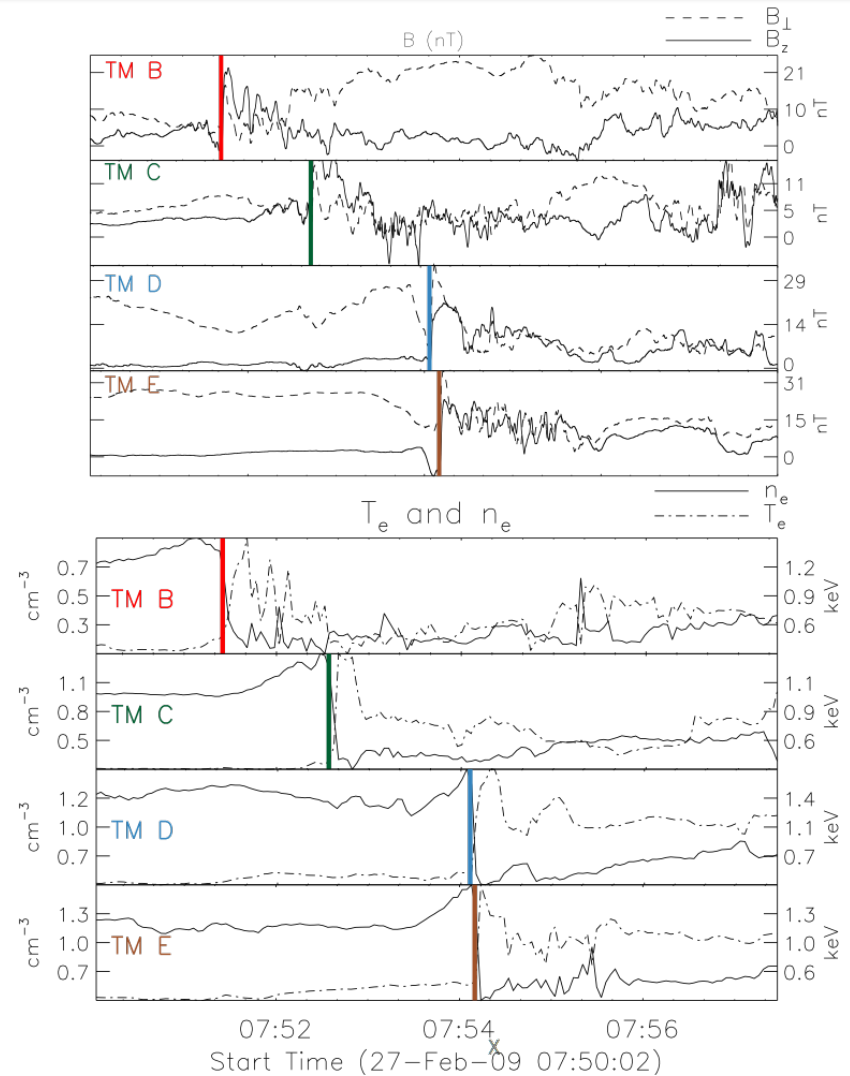
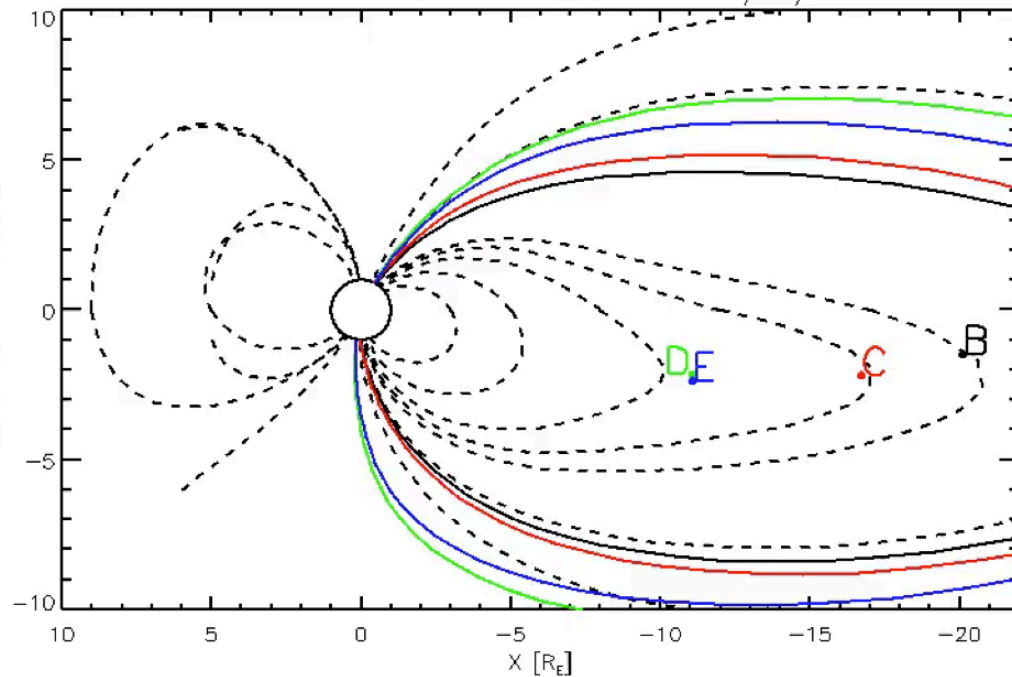


Substantial density drop following the dipolarization event!

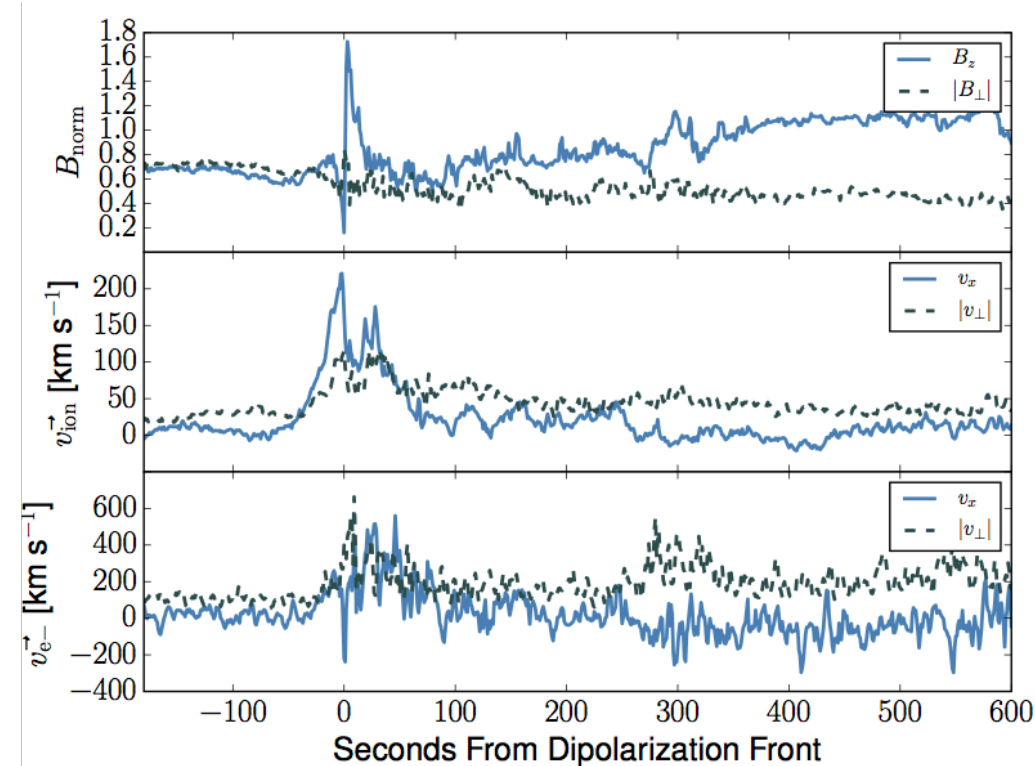
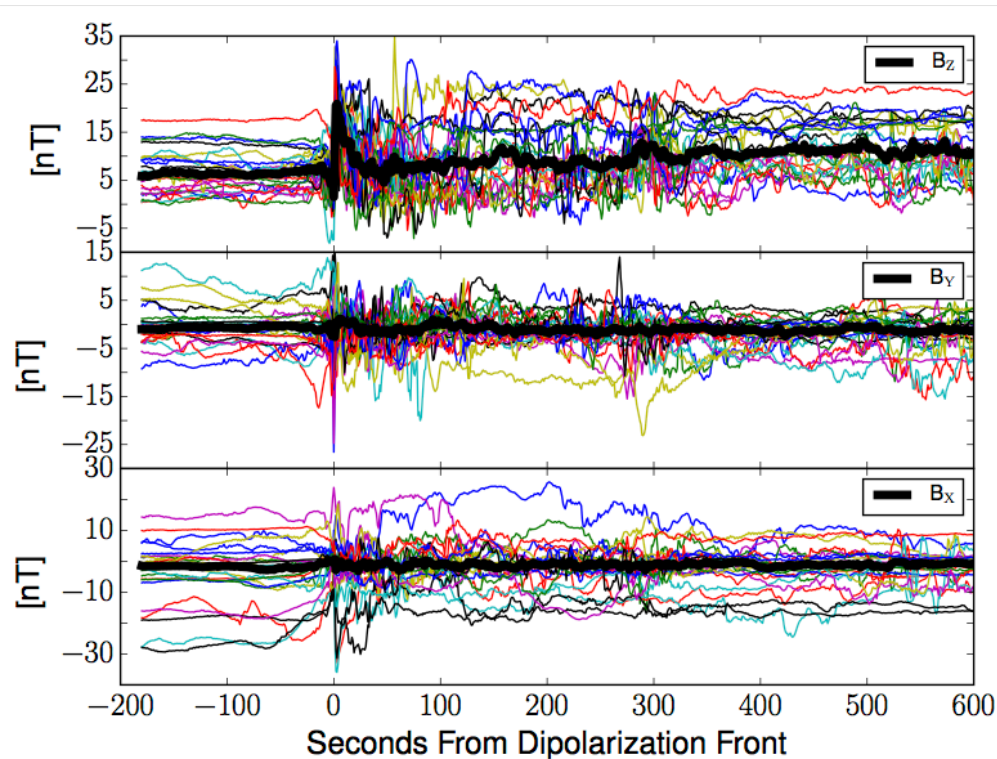
Using THEMIS for comparison to *in situ* data



Model t01 and Location of THEMIS Probes 2009/02/27 07:44:00

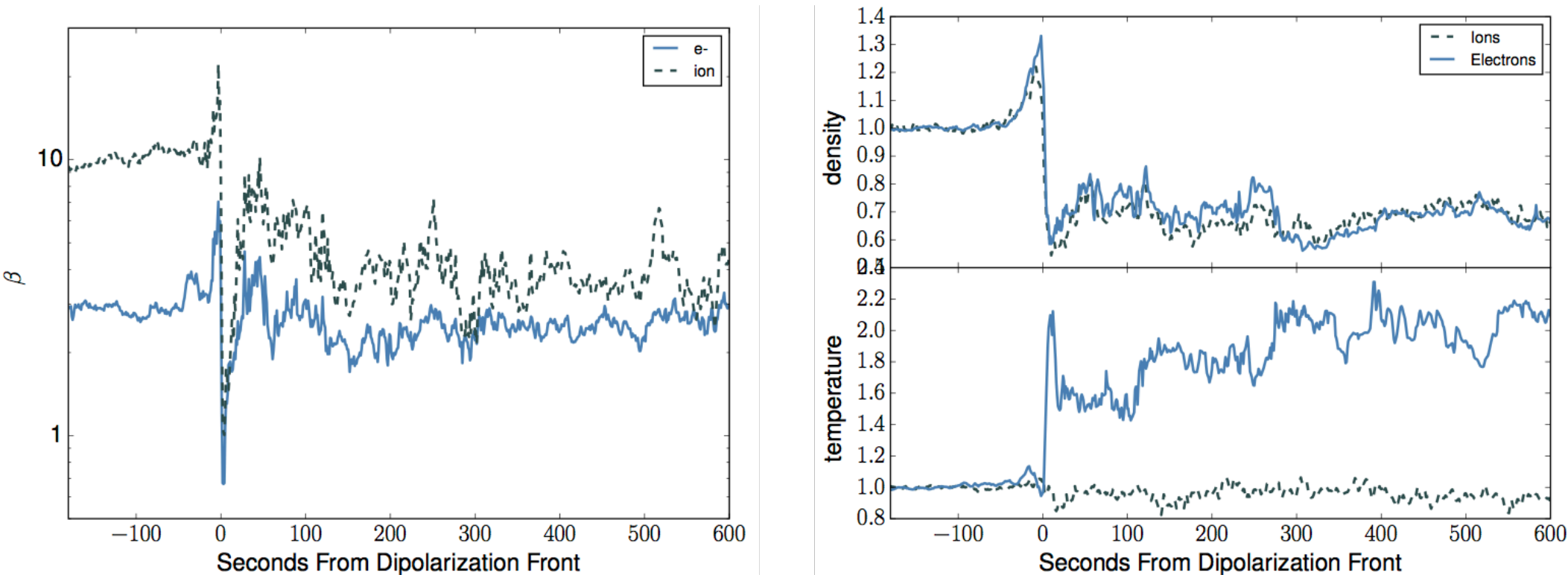


Using THEMIS for comparison to *in situ* data



Combination of 6 events observed by up to 5 THEMIS spacecraft
(listed in Runov et al., 2011)

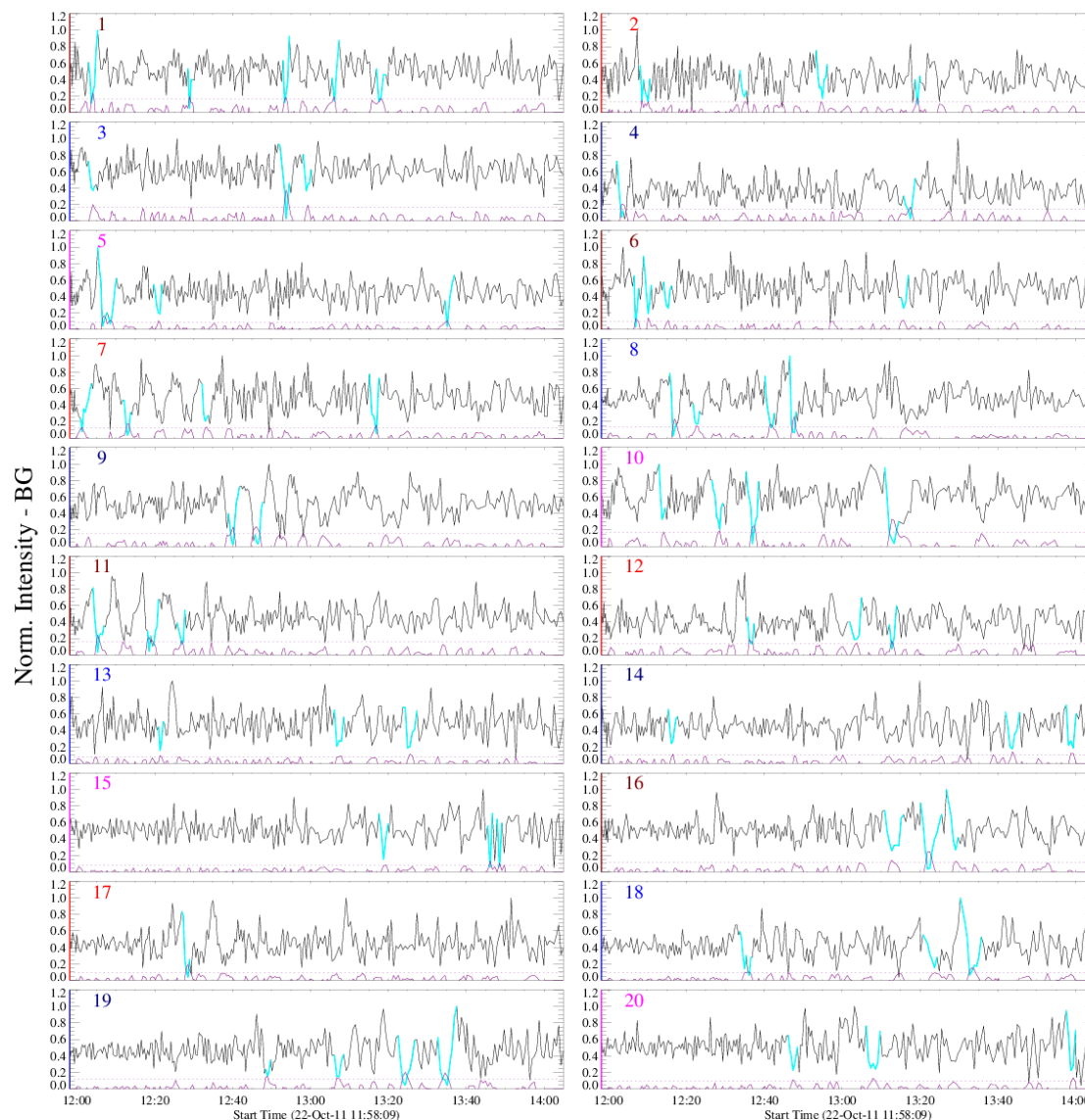
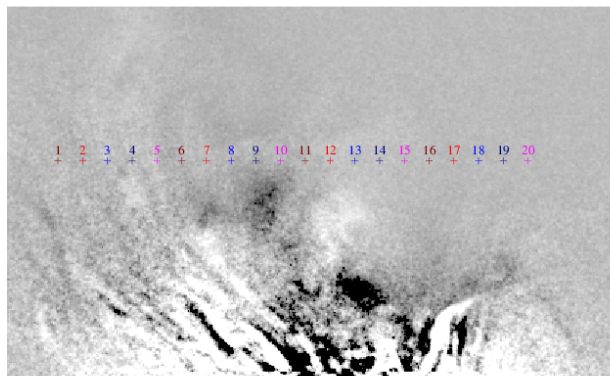
Using THEMIS for comparison to *in situ* data



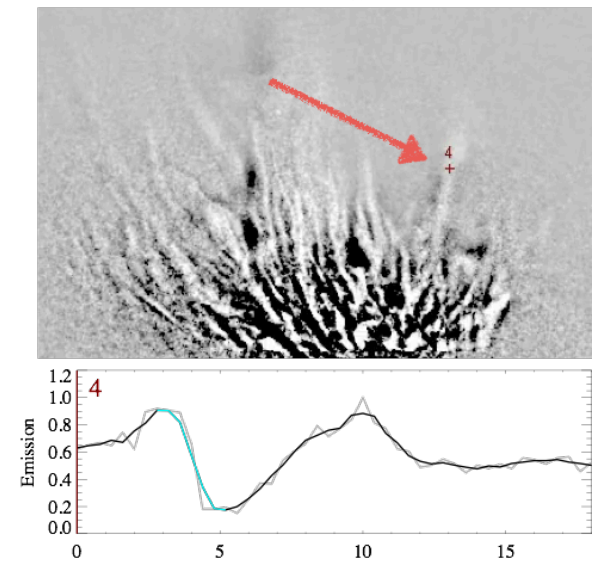
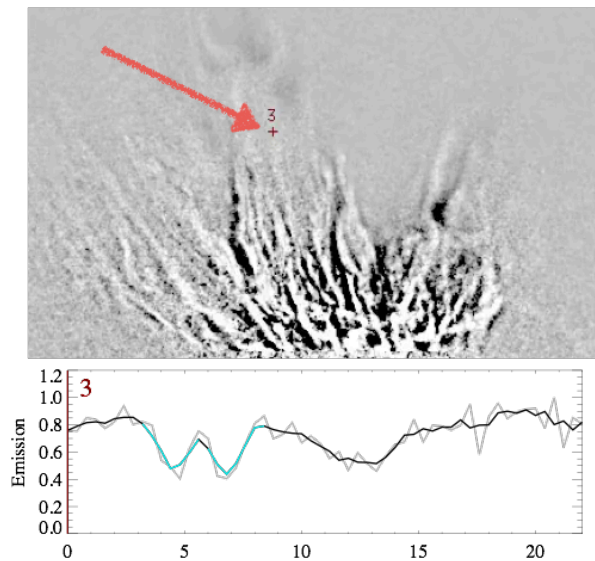
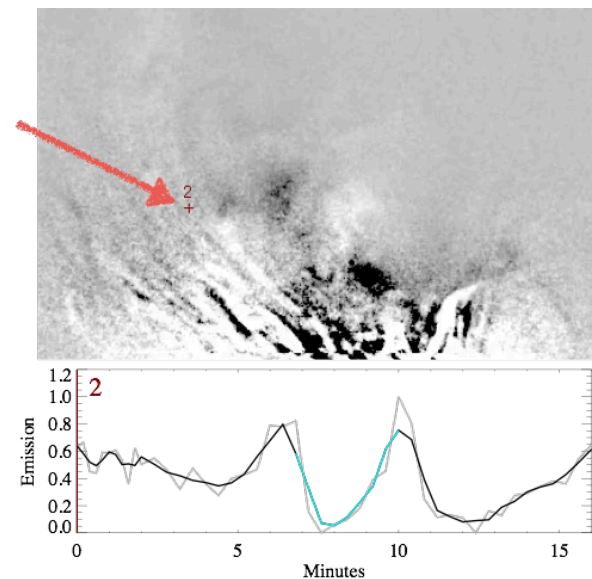
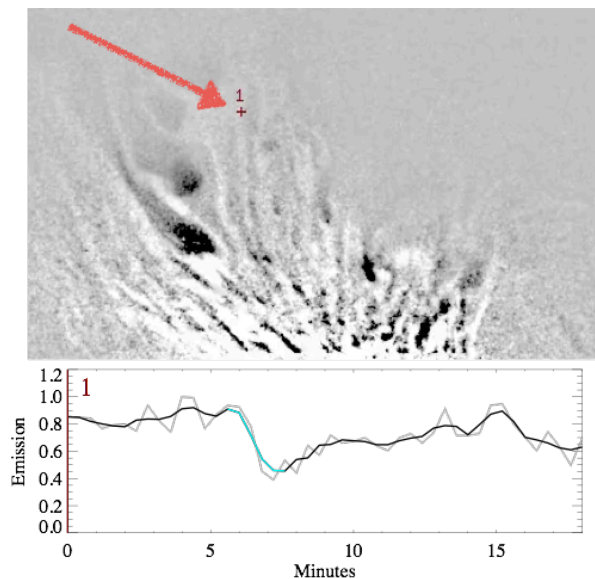
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Mimicking *in situ* data sets in the corona

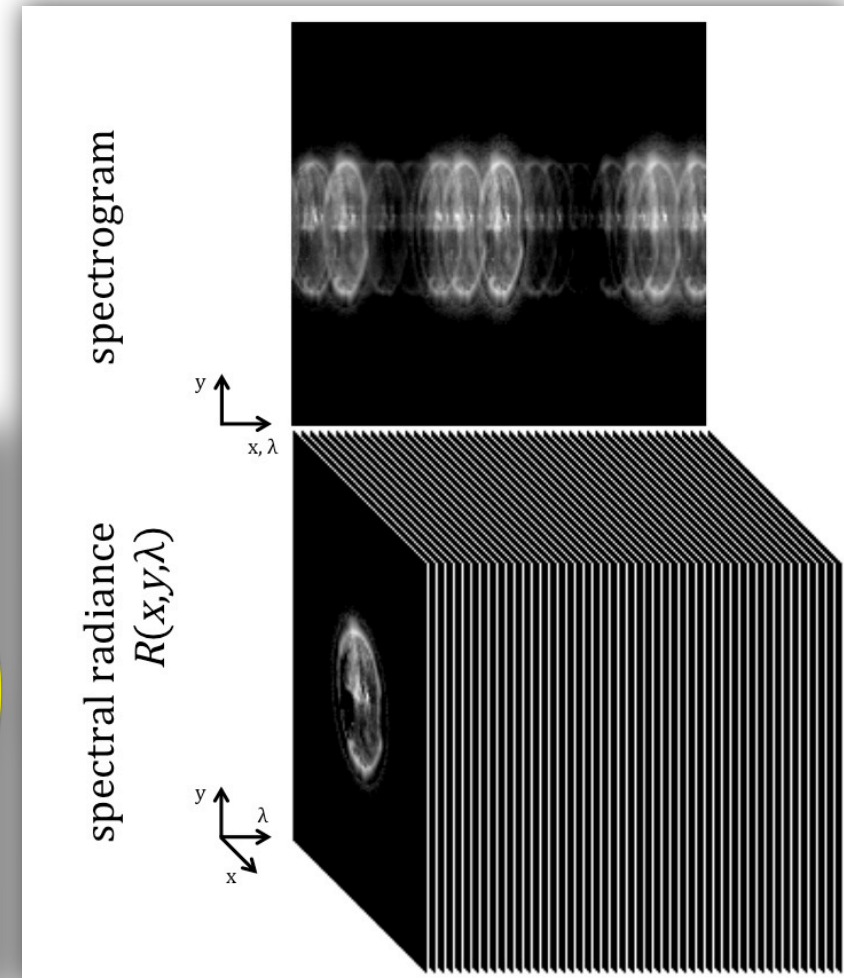
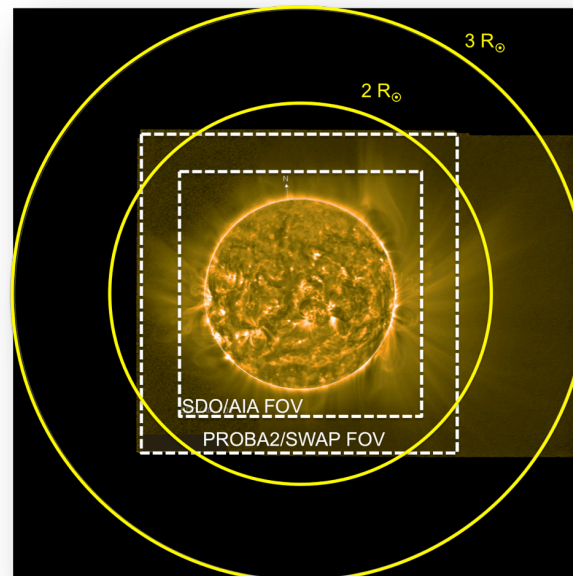


Mimicking *in situ* data sets in the corona



Looking forward

- Need for instrumentation to fill the gap in observing the transition corona
 - Important to be in single wavelength
 - Proposed for the International Space Station as a MoO; awarded Technology Development funding....
[COSIE]



Take aways

- *In situ* magnetotail data being used to inform remote sensing coronal data (and eventually vice versa...)
- Continuation of shrinking loops can impart energy into the current sheet **long** after the eruption and **high** into the corona.
- COSIE instrument would immensely add to our transitional corona knowledge

